

TIME-DEPENDENT PERTURBATION THEORY [SEC. 2]

If we can write:

$$\hat{H}(r, t) = H_0(r) + \lambda V(r, t) \quad (2.1)$$

λ is small

If the Hamiltonian is the sum of a stationary part (\hat{H}_0) and a small time-dependent perturbation, $\lambda V(r, t)$, then we can seek analytical solutions.

Assume we know the solutions of \hat{H}_0 :

$$\hat{H}_0 \psi_n^{(0)}(r) = E_n^{(0)} \psi_n^{(0)}(r) \quad (2.2)$$

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unperturbed ψ functions and eigenvalues

For the unperturbed case, the general solution of the TDSE:

$$\hat{H}_0 \Psi(r, t) = i\hbar \frac{d\Psi(r, t)}{dt} \quad (2.3)$$

(as seen in Sec 1.2) reduces to the time-indep. conservative form:

$$\Psi(r, t) = \sum_n C_n^{(0)} \psi_n^{(0)} e^{-iE_n^{(0)}t/\hbar} \quad (2.4)$$

the unperturbed prob amplitudes $C_n^{(0)} = \langle \psi_n^{(0)} | \Psi(r, t=0) \rangle$

are constant in time.

[2.1] DIRAC'S METHOD OF VARIATION OF CONSTANTS.

If $\lambda V(r, t) \neq 0$, the energy is not conserved

NB unperturbed case is conservative since $\langle E \rangle = \langle \hat{H}_0 \rangle$
 $\langle \Psi | \hat{H}_0 | \Psi(t) \rangle = \sum_n |C_n^{(0)}|^2 E_n$

$\langle E \rangle$ is independent of time.

But even with the perturbation, we express the new solutions in terms of the unpert. functions

$\Psi_n^{(0)}(r)$, but now:

$$\Psi(r, t) = \sum_n C_n(t) \Psi_n^{(0)} e^{-i E_n^{(0)} t / \hbar} \quad (2.5)$$

... We allow the coefficients to vary with time.

We still need normalisation so $\sum_n |C_n(t)|^2 = 1$ (2.6)

From orthonormality of $\Psi_n^{(0)}(r)$

$$\langle \Psi_n^{(0)} | \Psi(r, t) \rangle = C_n(t) e^{-i E_n^{(0)} t / \hbar} \quad (2.7)$$

To calculate the $C_n(t)$, substitute eq (2.5)

into the T.D.S.E, ($i\hbar \partial \Psi / \partial t = H \Psi$) to obtain:

$$i\hbar \sum_n \dot{C}_n \Psi_n^{(0)} e^{-i E_n^{(0)} t / \hbar} + \sum_n C_n(t) \Psi_n^{(0)} E_n^{(0)} e^{-i E_n^{(0)} t / \hbar} = \sum_n (H_0 + \lambda V) C_n(t) \Psi_n^{(0)} e^{-i E_n^{(0)} t / \hbar} \quad (2.8)$$

but, $H_0 \Psi_n^{(0)} = E_n^{(0)} \Psi_n^{(0)}$

As the 2nd and 3rd terms cancel, giving:

$$i\hbar \sum \dot{c}_n \Psi_n^{(0)} e^{-iE_n^0 t/\hbar} = \lambda \sum_n V(r,t) c_n \Psi_n^{(0)} e^{-iE_n^0 t/\hbar} \quad (2.9)$$

Take the scalar product of (2.9) (both sides) with $\langle \Psi_k^0 |$ i.e. $\Psi_k^{0*}(r)$ to obtain:

$$i\hbar \dot{c}_k e^{-iE_k^0 t/\hbar} = \lambda \sum_n V_{kn}(t) c_n(t) e^{-iE_n^0 t/\hbar} \quad (2.10)$$

since $\langle \Psi_k^0 | \Psi_n^0 \rangle = \delta_{kn}$.

$$V_{kn}(t) = \langle \Psi_k^0 | V(r,t) | \Psi_n^0 \rangle \quad (2.11)$$

is 'THE MATRIX ELEMENT OF THE PERTURBATION'.

Hence,

$$\frac{dc_k(t)}{dt} = \frac{\lambda}{i\hbar} \sum_n V_{kn}(t) e^{i\omega_{kn}t} c_n(t) \quad (2.12a)$$

$\omega_{kn} = (E_k^0 - E_n^0) / \hbar$ = the BOHR ANGULAR FREQUENCY.

(2.12) represents a set of coupled linear differential equations. We can write it in compact matrix form:

$$\dot{\vec{C}} = \frac{\lambda}{i\hbar} V \vec{C} \quad (2.12b)$$

$$\begin{bmatrix} \dot{c}_1 \\ \dot{c}_2 \\ \vdots \\ \dot{c}_N \end{bmatrix} = \begin{bmatrix} V_{11} & V_{12}(t)e^{i\omega_{12}t} & V_{13}(t)e^{i\omega_{13}t} & \dots \\ V_{21}e^{-i\omega_{12}t} & V_{22} & & \\ V_{31}e^{-i\omega_{13}t} & & V_{33} & \\ & & & \ddots \\ & & & & V_{NN} \end{bmatrix} \begin{bmatrix} c_1 \\ c_2 \\ \vdots \\ c_N \end{bmatrix}$$

These equations can be solved numerically, without further assumption.

Perturbation theory means the further assumption that λ is small.

[2.2] PERTURBATION EXPANSION

Assume perturbation is weak and expand the coefficients C_n in powers of λ :

$$C_n = C_n^{(0)} + \lambda C_n^{(1)} + \lambda^2 C_n^{(2)} \dots = \sum_{j=0}^{\infty} \lambda^j C_n^{(j)}$$

Subs into (2.12): (2.13)

$$\sum_j \lambda^j \dot{C}_k^{(j)} = \frac{\lambda}{i\hbar} \sum_n V_{kn}(t) e^{i\omega_{kn}t} \sum_j \lambda^j C_n^{(j)}$$

(2.14)

Equate powers of λ :

λ^0 (0-th order) $C_k^{(0)} = 0$ (0-th order is time-indep.) 2.15(a)

λ^1 (1st order) $\dot{C}_k^{(1)} = \frac{1}{i\hbar} \sum_n V_{kn} e^{i\omega_{kn}t} C_n^{(0)}$ 2.15(b)

λ^2 (2nd order) $\dot{C}_k^{(2)} = \frac{1}{i\hbar} \sum_n V_{kn} e^{i\omega_{kn}t} C_n^{(1)}$ 2.15(c)

(2.15)

In principle, we can integrate successively to any order.

1ST ORDER PERTURBATION CORRECTION

Assume that initially (at $t=t_0$) the system is in a particular unperturbed state $\psi_i^{(0)}$

Then $C_k^{(0)} = \delta_{ki}$ or $\delta(k-i)$
DISCRETE CONTINUOUS

Hence, in (2.15 b) we have:

$$\dot{C}_k^{(1)} = \frac{1}{i\hbar} V_{ki} e^{i\omega_{ki}t}$$

Hence, we have a correction to the initial state $k=i$,

$$C_i^{(1)}(t) = \frac{1}{i\hbar} \int_{t_0}^t V_{ii}(t') dt' \quad (2.16)$$

and a transfer of amplitude into other eigenstates $i \neq k$

$$C_k^{(1)}(t) = \frac{1}{i\hbar} \int_{t_0}^t V_{ki}(t') e^{i\omega_{ki}t'} dt' \quad (2.17)$$

The constant of integration is chosen so that:

$$C_k^{(1)}(t) = \delta_{ki} \text{ at } t=t_0$$

TO 1ST ORDER :

$$C_k = \delta_{ki} + \lambda C_k^{(1)} \quad (2.18)$$

There has been a transfer of population to other eigenstates. The probability of finding the system, initially in state i , to now be in state $k \neq i$

$$P_{ki}^{(1)}(t) = |C_k(t)|^2 = |\lambda C_k^{(1)}|^2$$

$$= \frac{1}{\hbar^2} \left| \int_{t_0}^t \lambda V_{ki} e^{i\omega_{ki}t'} dt' \right|^2$$

(2.19)

While the change to the initial state i , in the perturbative regime is:

$$C_i(t) \approx 1 + \lambda C_i^{(1)} \approx 1 + \frac{\lambda}{i\hbar} \int_{t_0}^t V_{ii}(t') dt'$$

if we can approximate this by

$$\approx \exp - \frac{i}{\hbar} \int_{t_0}^t \lambda V_{ii}(t') dt' \quad (2.20)$$

then

$$|C_i(t)|^2 \approx 1 \quad \text{and the main}$$

effect of the perturbation is to change the phase of the initial quantum state

2ND ORDER AMPLITUDES

Sometimes we need to go to higher order, eg 1) if 1st-order correction vanishes b) if λ not small enough, so $C_n = C_n^{(0)} + \lambda C_n^{(1)} + \lambda^2 C_n^{(2)} \dots$ falls off slowly with increase in order.

We go back to 2.15(c)

$$\dot{C}_k^{(2)} = \frac{1}{i\hbar} \sum_n V_{kn}(t) e^{i\omega_{kn}t} C_n^{(1)}(t)$$

where, $C_n^{(1)}(t) = \frac{1}{i\hbar} \int_{t_0}^t V_{ni}(t') e^{i\omega_{ni}t'} dt'$ where $C_n^{(0)} = \delta_{ni}$

$$C_k^{(2)}(t) = \frac{1}{i\hbar} \int_{t_0}^t \sum_n V_{kn}(t'') e^{i\omega_{kn}t''} C_n^{(1)}(t'') dt''$$

now subs. the $C_n^{(1)}(t'')$

$$= \frac{1}{i\hbar} \int_{t_0}^t \sum_n V_{kn}(t'') e^{i\omega_{kn}t''} \left[\frac{1}{i\hbar} \int_{t_0}^{t''} V_{ni}(t') e^{i\omega_{ni}t'} dt' dt'' \right]$$

$$= -\frac{1}{\hbar^2} \int_{t_0}^t dt'' \int_{t_0}^{t''} dt' \sum_n V_{kn}(t'') V_{ni}(t') e^{i\omega_{kn}t''} e^{i\omega_{ni}t'} \tag{2.21}$$

So, even if matrix element $V_{ki} = 0$ (no direct coupling from initial to final state) we can go $i \rightarrow k$ via intermediate states n

[2.3] A CONSTANT PERTURBATION

A simple example is given by a perturbation $V(r)$ which is switched on at $t = t_0$ and off again at t but is constant in between



For convenience, choose $t_0 = 0$ and $\lambda = 1$ so

$$\hat{H} = \begin{cases} H_0 & t' < 0 \\ H_0 + V(r) & 0 < t' < t \\ H_0 & t' > t \end{cases}$$

If our system is in state $\psi_i^{(0)}(r)$ at $t' < 0$, at a later time $t' > t$ (ie after switch-off) the amplitude for state k is

$$C_k^{(1)} = \frac{1}{i\hbar} \int_0^t V_{ki} e^{i\omega_{ki}t'} dt' \quad k \neq i \quad (2.22)$$

$V_{ki} = \int \psi_k^*(r) V(r) \psi_i(r) dr$ is just a number...

$$\begin{aligned} \text{so } C_{ki}^{(1)} &= \frac{V_{ki}}{i\hbar} \int_0^t e^{i\omega_{ki}t'} dt' \\ &= \frac{V_{ki}}{\hbar\omega_{ki}} [1 - e^{i\omega_{ki}t}] \quad (2.23) \end{aligned}$$

The corresponding probability, to 1st order

$$P_{ki}^{(1)}(t) = |C_k^{(1)}(t)|^2$$

$$= \frac{1}{\hbar^2} |V_{ki}|^2 \frac{(1 - e^{i\omega_{ki}t})(1 - e^{-i\omega_{ki}t})}{\omega_{ki}^2}$$

$$P_{ik} = \frac{4}{\hbar^2} |V_{ki}|^2 \frac{\sin^2 \omega_{ki} \cdot t/2}{\omega_{ki}^2} \quad (2.24)$$

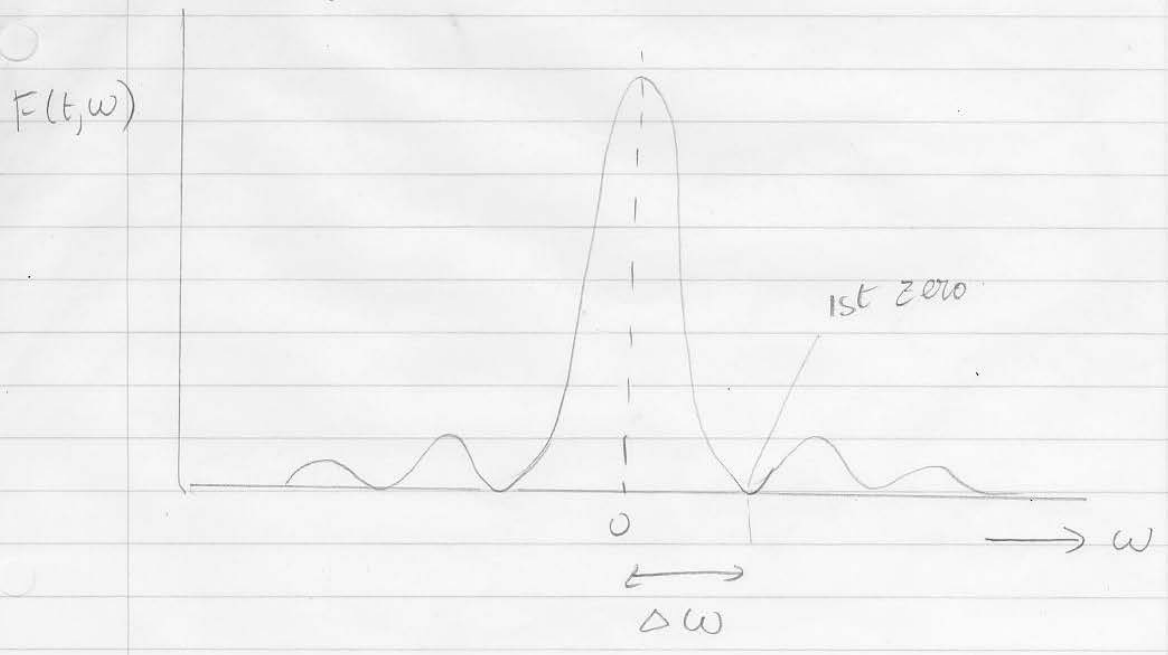
so,
we write

$$P_{ik}^{(1)}(t) = \frac{2}{\hbar^2} |V_{ki}|^2 F(t, \omega_{ki}) \quad (2.25)$$

the function $F(t, \omega) = 2 \frac{\sin^2 \omega t/2}{\omega^2}$

$$F(t, \omega) = t^2/2 \frac{\sin^2 \omega t/2}{(\omega t/2)^2} = t^2/2 \left(\text{sinc} \frac{\omega t}{2} \right)^2 \quad (2.26)$$

F is sharply peaked about $\omega = 0$



- 1) width of central peak $\Delta\omega = 2\pi/t$ (1st zero of $\sin \omega t$)
- 2) height of the peak = $t^2/2$
- 3) with increasing time, peak gets taller and thinner.

in fact, $\lim_{t \rightarrow \infty} F(t, \omega) = \pi t \delta(\omega)$ (see I.4) (2.27)

So, for short times (short lived perturbation), P_{ik} the probability, allows almost any energy jump.
 For long times, jumps with $\omega_{ik} = E_i - E_k = 0$ are increasingly favoured.

This is a version of the energy-time uncertainty relation

$t \equiv \Delta t \rightarrow$ length of perturbation

$$\Delta\omega = \frac{2\pi}{t} = \frac{2\pi}{\Delta t} \quad \Delta E = \hbar \Delta\omega = \frac{2\pi \hbar}{\Delta t}$$

so, $\Rightarrow \Delta E \Delta t \gtrsim \hbar$ (2.28)

From eq (2.25), the probability at time t

$$P_{ik}^{(1)}(t) = \frac{2}{\hbar^2} |V_{ki}|^2 F(t, \omega_{ki})$$

$$\text{Now, } \delta(\omega_{ki}) = \hbar \delta(E_i - E_k) \quad \Downarrow \quad \pi t \delta(\omega_{ki})$$

(From the property of the Dirac delta function) ⊙

$$\delta(ax) = \frac{1}{|a|} \delta(x) \quad (\text{eq I.4d})$$

$$\text{and } \hbar \omega_{ik} = E_i - E_k$$

So

$$P_{ik}^{(1)}(t) = \frac{2\pi t}{\hbar} |V_{ki}|^2 \delta(E_k - E_i) \quad (2.29)$$

We can define a transition probability per unit time,
a transition rate

$$\Gamma = \frac{2\pi}{\hbar} |V_{ki}|^2 \delta(E_k - E_i) \quad (2.30)$$

'Fermi's golden rule'